

EFT and GR - Saalberg 4

9/13/13
872972013

Old view : GR + QM incompatible

New view : GR + QM = EFT

EFT have limits

Quantum + Gravity completion UV

↑ Many interesting questions

Progress

Graviton-graviton scattering

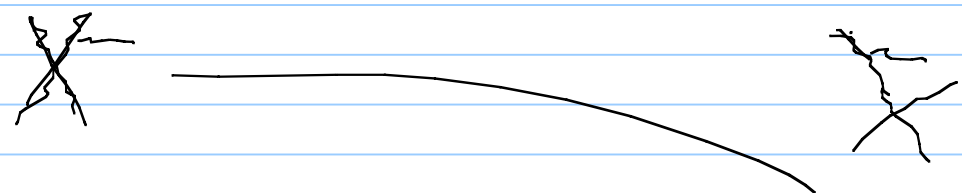


Dimitar
Noveckiy

$$\begin{aligned}
 \mathcal{A}(++;++) &= \frac{i \kappa^2 s^3}{4 tu} \left(1 + \frac{\kappa^2 s t u}{4(4\pi)^{2-\epsilon}} \frac{\Gamma^2(1-\epsilon)\Gamma(1+\epsilon)}{\Gamma(1-2\epsilon)} \times \right. \\
 &\times \left[\frac{2}{\epsilon} \left(\frac{\ln(-u)}{st} + \frac{\ln(-t)}{su} + \frac{\ln(-s)}{tu} \right) + \frac{1}{s^2} f\left(\frac{-t}{s}, \frac{-u}{s}\right) \right. \\
 &\left. \left. + 2 \left(\frac{\ln(-u)\ln(-s)}{su} + \frac{\ln(-t)\ln(-s)}{tu} + \frac{\ln(-t)\ln(-s)}{ts} \right) \right] \right)
 \end{aligned}$$

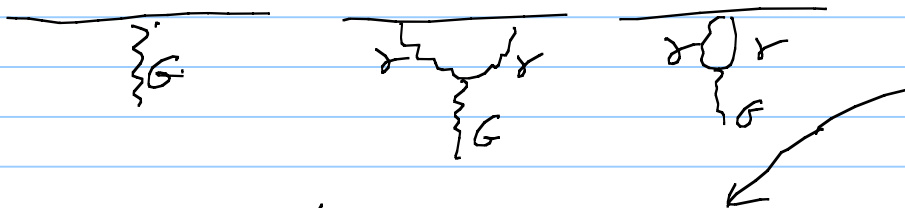
IR
—

no C_1, C_2 - pure gravity 1-loop finite



$$d\sigma(GG \rightarrow GG) + d\sigma(GG \rightarrow GG\gamma) = \text{finite}$$

Corrections to Reissner Nordstrom + Kerr Newman Metric (M, Q, S)
 - QED



$$T_{\mu\nu}(E) = \begin{pmatrix} \frac{1}{2} \mathbf{F} \cdot \mathbf{F} & \\ & -\mathbf{F} \cdot \mathbf{F} \end{pmatrix}$$

$$g_{00} = 1 - \frac{2Gm}{r} + \frac{G\alpha}{r^2} - \frac{2}{3} \frac{G\alpha h}{\pi m r^3}$$

$$g_{ij} = \left(\dots \dots \frac{2}{\pi} \frac{G\alpha h}{m r^3} \right) (\vec{S} \times \vec{S})_i$$

Limits of EFT

$$M = G E^2 \left(1 + G E^2 + \beta E^2 \ln E \right)$$

$\leftarrow \mathcal{O}(1) = M_P^2 = E^2$

Low E issues

R small

\leftarrow potential long issues

M_P



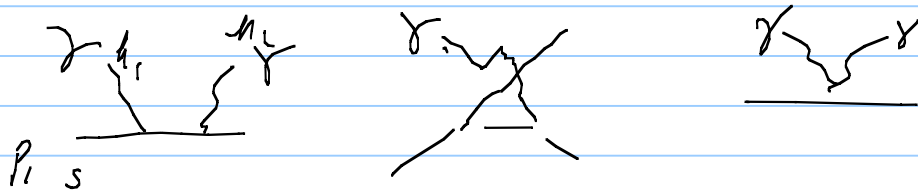
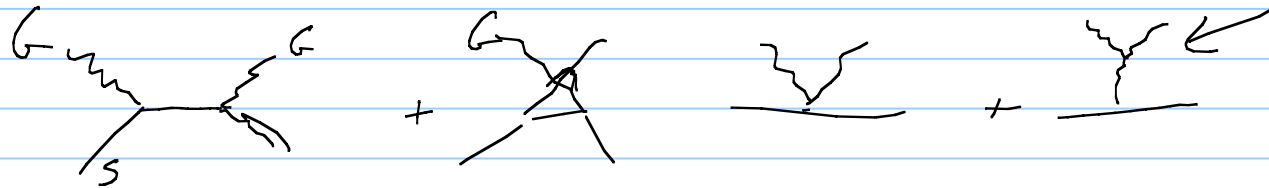
You

\leftarrow R small, Integrated curvature $R y^2$

$$g_{\mu\nu} = \eta_{\mu\nu} + \underline{R_{\mu\nu\alpha\beta}} \underline{y^\alpha} \underline{y^\beta} + \dots$$

Gravity = "Square of Gauge Theory"

KLT $\circ \cup$



$$M_{G_s}(\quad) = \frac{\kappa^2}{2s^2} \frac{p_1 \cdot k_1 p_1 \cdot k_2}{k_1 \cdot k_2} A_{\gamma_s}(\quad) A_{s_0}(\quad)$$

$$E_{\mu\nu}(+2) = E_{\mu}(+1) E_{\nu}(+1)$$

Helicity amplitude

$\bar{u}(h_2) \gamma u(h_1)$

$$A_0 = \frac{m^2 s [h_1 h_2]^2}{4(p_1 \cdot h_1)(p_1 \cdot h_2)}$$

L. Dixon TASI

R. K. Ellis

$$M_G = \frac{k^2}{16 k_1 \cdot k_2} \frac{[k_1 k_2]^2}{p_1 \cdot h_1 p_2 \cdot h_2}$$

On shell methods

Real parts \sim Imag parts

\leftarrow on shell states

- } - } \leftarrow on shell

1) Dispersion relation

$$V(q^2) = \frac{i}{\pi} \int dt \frac{\rho(t)}{t - q^2}$$

\leftarrow on shell states only

2) Matching to box, triangle + bubble diagrams

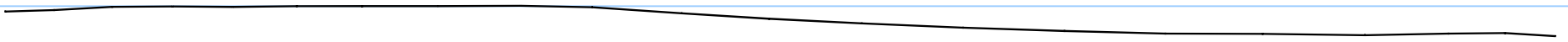
II X O

- Passarino-Veltman - reduced to \leftarrow

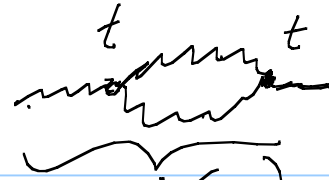
- identifiable cuts \leftarrow

$$M = N_B \text{Box}(p_1, p_2, k_1, k_2) + N_T \text{Triangle}(\quad) + N_{\text{Bub}} \text{Bubble (Born)}$$

Bjerrum Bohr, Vanhove, JFD



Cosmology



B. El Menoufi

$$\left(\frac{\dot{a}}{a}\right)_t^2 = \frac{8\pi G}{3} T_{\mu\nu}(t) + G t \left(\frac{\dot{a}}{a}\right)_t^2 \int_{-\infty}^t \frac{dt'}{t-t'}$$



Memory of past expansion

QFT + Cosmology

Scattering

$$S_{f_1} = \langle \psi_f(+\infty) | U_I(+\infty, -\infty) | \psi_i(-\infty) \rangle$$

$$\mathcal{M} = \langle 0 | T(O(t) U_I(+\infty, -\infty)) | 0 \rangle$$

"in-m", "closed time path", "Schwinger Keldysh"

$$\mathcal{M} = \langle \psi(-\infty) | U^\dagger(-\infty, t) T_{\mu\nu}(t) U(t, -\infty) | \psi(-\infty) \rangle$$

$$= \langle \psi(-\infty) |$$

$$\underbrace{U^\dagger(-\infty, t) U^\dagger(t, \infty)}_{U^\dagger(-\infty, \infty)} \underbrace{U(\infty, t) T_{\mu\nu}(t) U(t, -\infty)}_{T(T_{\mu\nu}(t) U(\infty, -\infty))} | \psi(-\infty) \rangle$$

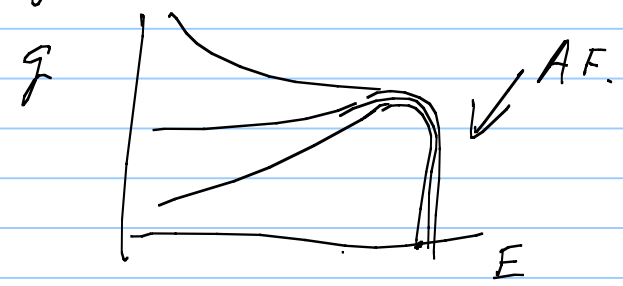
abnt. time ordered!
usual

Velkovisky

Running couplings + gravity

Robinson + Wilczek $\beta(g, E) = b g^3 + c g K^2 E^2$

- 1) RW wrong g does not run
- dim reg
- 2) RW right g does run
- cutoff reg



EFT to rescue

$$M_i = a_i e^2 + e^2 \left[b_i k^2 q^2 + c_i k^2 q^2 \ln q^2 \right]$$

← one process
"

RW renormalize at $q^2 = E^2$

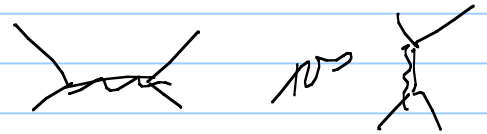
$$= a_i e^2 \left[1 + \frac{b_i}{a_i} k^2 E^2 \right] + e^2 b_i k^2 \underbrace{(q^2 - E^2)}$$

$$= a_i e^2(E)$$

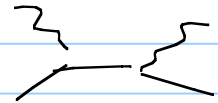
But:

1) Kinematic $q^2 = +E^2$, $q^2 = -E^2$

$e^2(E) \downarrow$ or \uparrow ← not universal



2) Not universal a_i, b_i, c_i depend on process



Do calculation \Rightarrow mo' useful or universal def of $e^2(F^2)$

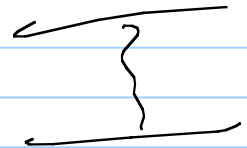
Cutoff: $e A \rightarrow A$

$$\mathcal{L} = \frac{1}{4e_0^2} F^2 + \bar{\Psi}(i\not{D} - m)\Psi$$

\leftarrow no elect charge

one loop

$$= \frac{1 + c k^2 \Lambda^2}{4e_0^2} F^2 - \dots$$



$$\beta(\Lambda) = \Lambda \frac{\partial}{\partial \Lambda} \left(\frac{e^2}{1 + c k^2 \Lambda^2} \right) \neq 0 \quad \Rightarrow \text{running } e$$

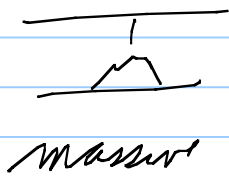
$$\frac{l^2}{4\pi} = \frac{l_0^2}{(1+k)^2} \frac{1}{4\pi} = \frac{1}{137}$$

Running e with gravity is no a good concept

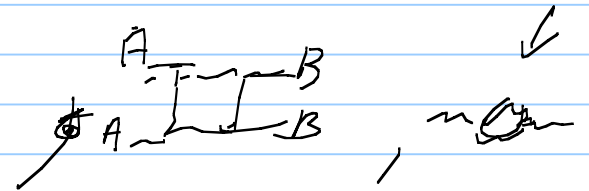
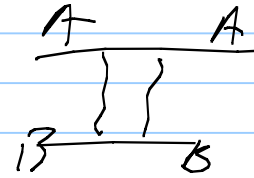
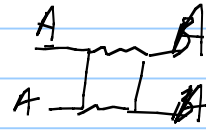
Running $G(E)$?

- Λ cutoff wrong

Calculations



graviton
~~loop~~



6 def of $G(E) \rightarrow$ Wildly different \rightarrow sign
 \rightarrow magnitude

- * 1) $g^2 = \pm E^2$ - ambiguity \in
 - 2) Non universality
- } $G(E)$ in Lorentzian spaces
 is not useful or universal.