

Polaron in the weak coupling limit.

Consider a semiconductor. Electrons in a conduction band represent a diluted gas. When concentration is small, electron-electron interactions can be neglected.

At the same time, ^{an injected} ~~as single~~ electrons, while moving in the a crystallized lattice, it polarizes its surrounding medium and leads to a deformation of the lattice. These deformations are nothing but phonons. Such an electron, which is surrounded by a cloud of phonons, is called polaron.

$$\text{Polaronic dispersion is: } \varepsilon(\vec{p}) = \varepsilon_0 + \frac{\vec{p}^2}{2m^*}$$

$\varepsilon_0 m^*$ - effective mass, which we will calculate.

To be more specific, we let us calculate the lowest-order electron self-energy in the lowest order in e-ph interactions. It is given by this diagram



where $G(\varepsilon, \vec{p}) = \frac{1}{\varepsilon - \frac{\vec{p}^2}{2m} + i\delta}$,

$\delta > 0 \Leftrightarrow$ electron tunnels into the semiconductor from outside, and ^{where} there was in the conductor zone ^{there were} no electrons before.

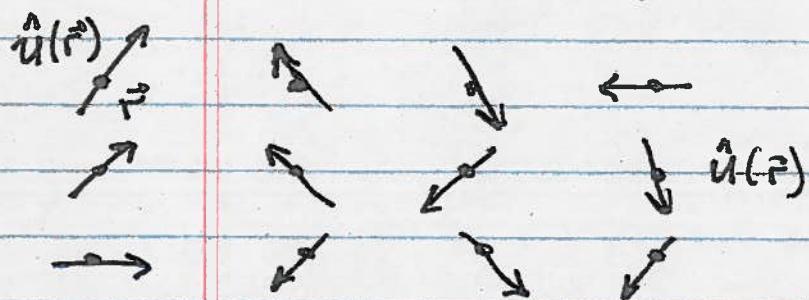
We calculate $Z(\varepsilon, \vec{p})$ in the vicinity of the mass-shell defined by a condition $\varepsilon = \frac{\vec{p}^2}{2m}$, and at $|\vec{p}| \ll mc$.

Electron-phonon interactions

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Let $\hat{u}(\vec{r}, t)$ be the operator associated with

~~displacements of atoms on a lattice~~



$\hat{u}(\vec{r})$ forms dipole moments and induces polarization density

$$\vec{P}(\vec{r}) = e \cdot \frac{N}{V} \cdot \hat{u}(\vec{r})$$

N - total # of atoms

V - total volume.

Generally, ~~the~~ electric displacement field (induced field) is given by

$$\vec{D} = \epsilon_0 \cdot \vec{E} + 4\pi \vec{P}(\vec{r}) = \epsilon \cdot \vec{E}$$

ϵ_0 (vacuum) dielectric constant

From Maxwell equations we have ($\vec{E} = 0$)

$$\operatorname{div} \vec{D} = 4\pi \rho \Rightarrow \boxed{\operatorname{div} \vec{P}(\vec{r}) = \rho_c} \approx \operatorname{div} \hat{u}(\vec{r})$$

→ where ρ_c is the charge density induced by the displacement field, $\hat{u}(\vec{r})$.

Deformation potential $V(\vec{r}) = g \cdot \hat{\phi}(\vec{r}) \approx c \sqrt{\rho} \operatorname{div} \hat{u}(\vec{r}) \approx \rho_c$.

Interaction Hamiltonian is a product of densities given by a product of densities:

$$H_{\text{int}} \approx \int k(\vec{r}) K(\vec{r} - \vec{r}') \text{div} \vec{P}(\vec{r}') d^3 r' d^3 r$$

electron density electric charge density. - 5-

where K is an interaction function.

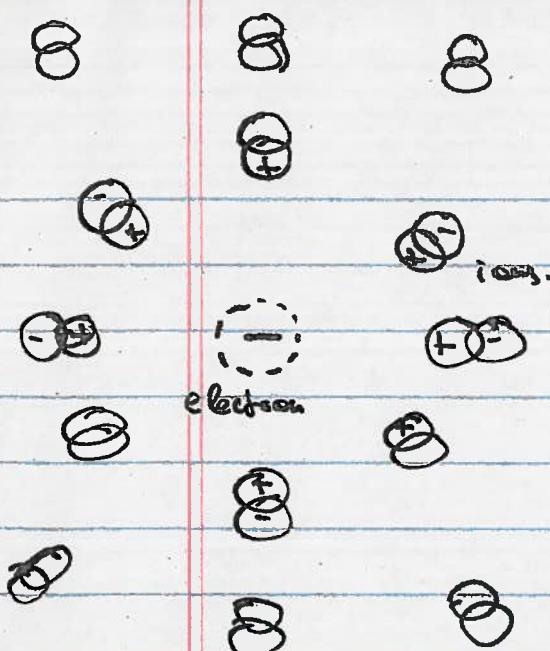
$$K(\vec{r} - \vec{r}') = \begin{cases} \frac{1}{|\vec{r} - \vec{r}'|}, & |\vec{r} - \vec{r}'| \ll a \\ 0, & |\vec{r} - \vec{r}'| \gg a \end{cases}$$

Approximation: $\underline{K(\vec{r} - \vec{r}') = \# \cdot \delta(\vec{r} - \vec{r}')} \Rightarrow$

$$\Rightarrow H_{\text{int}} = g \int \hat{\psi}^+(\vec{r}) \hat{\psi}(\vec{r}) \hat{\psi}(\vec{r}) d^3 r.$$

Electron- phonon interactions

1. Fröhlich Hamiltonian.
2. Polaron in a weak coupling limit.
3. Cherenkov radiation of sound.



Polaron: single electron

polarizes surrounding medium
and leads to a deformation
of the lattice.

These deformations are
phonons. As a result the
electronic effective mass
changes.

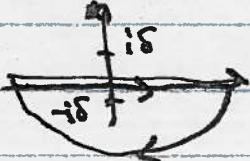
$$\frac{\epsilon, \vec{p}}{\epsilon - \omega, \vec{k}} = \sum (\epsilon, \vec{p}) = i g^2 \int G_0(\omega - \omega, \vec{p} - \vec{k}) D_0(\omega, \vec{k}) \frac{d^3 k}{(2\pi)^3} \frac{d\omega}{2\pi}$$

where $G_0(\epsilon, \vec{p}) = \frac{1}{\epsilon - p^2/m + i\delta}$ - electron Green's function

$D_0(\omega, \vec{k}) = \frac{c^2 k^2}{\omega^2 - c^2 k^2 + i\delta}$ - phonon Green's function.

Let us adopt the following sequence of integration:

1. over ω . $\int_{-\infty}^{\infty} d\omega \rightarrow$ choose



$$D_0(\omega, \vec{k}) = \frac{\omega_0(\vec{k})}{2} \left[\frac{1}{\omega - \omega_0(\vec{k}) + i\delta} - \frac{1}{\omega + \omega_0(\vec{k}) - i\delta} \right]$$

Let me first state the result, and then do the calculations. The result reads:

$$\Sigma(\varepsilon, \vec{p}) = \varepsilon_0 - d_1 \left(\varepsilon - \frac{\vec{p}^2}{2m} \right) - d_2 \frac{\vec{p}^2}{2m}$$

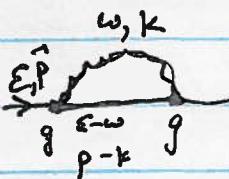
From here we can conclude that

- 1) d_2 - determines the mass renormalization: $\frac{m_f}{m} = f(d_2)$
Qualitatively this can be understood from the following reasoning....
- 2) d_1 - determines renormalization of fermion residue, \tilde{z} , which is not 1 anymore.

Proof:

$$G = \cancel{1} G_0 - \tilde{z} \int \left[\varepsilon - \frac{\vec{p}^2}{2m} + i\delta \right] - \varepsilon_0 - d_1 \left[\varepsilon - \frac{\vec{p}^2}{2m} \right] \cancel{d_2} \frac{\vec{p}^2}{2m} =$$

* calculation of the self-energy diagram:



$$= \Sigma(\varepsilon, \vec{p}) = ig^2 \int G_0(\varepsilon-\omega, \vec{p}-\vec{k}) D_0(\omega, \vec{k}) \frac{d^3 k}{(2\pi)^3} \frac{d\omega}{2\pi},$$

$$\text{where } G_0(\varepsilon, \vec{p}) = \frac{1}{\varepsilon - \frac{\vec{p}^2}{2m} + i\delta},$$

$$D_0(\omega, \vec{k}) = \frac{c^2 k^2}{\omega^2 - c^2 k^2 + i\delta}.$$

Let us perform integration over ω closing the contour in the lower half-plane:

$$\int_{-\infty}^{+\infty} \frac{1}{\tilde{\epsilon} - \omega + i0} \frac{e^2 k^2}{\omega^2 - c^2 k^2 + i0} \frac{d\omega}{2\pi} = \frac{i}{2} \frac{ck}{ck - \tilde{\epsilon} - i0}. \rightarrow \underline{\text{prove.}}$$

where $\tilde{\epsilon} = \epsilon - \frac{(p-k)^2}{2m}$, $k=|\vec{k}|$. Therefore we get

$$\Sigma(\epsilon, \vec{p}) = \frac{g^2}{2} \int \frac{ek}{\epsilon - ck - \frac{(p-k)^2}{2m} + i0} \frac{d^3k}{(2\pi)^3}$$

At the next step we make use of the following trick:

Integration over d^3k can be performed as follows.

Introduce $k=|\vec{k}|$, $q=|\vec{p}-\vec{k}|$, $x = \cos \hat{\vec{k}} \cdot \hat{\vec{p}} \Rightarrow$

$$\Rightarrow d^3k = 2\pi k^2 dk dx, \text{ while } q^2 = |\vec{p}-\vec{k}|^2 = p^2 + k^2 - 2pk \cos x \Rightarrow$$

$$\Rightarrow q dq = -pk dx.$$

We obtain that

$$\int f(1|\vec{k}|, 1|\vec{p}-\vec{k}|) \frac{d^3k}{(2\pi)^3} = \frac{1}{(2\pi)^2 \cdot p} \int_{|p-k|}^{|p+k|} k dk \int f(k, q) q dq.$$

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$$\Sigma(\varepsilon, \vec{p}) = ig^2 \int_{-\infty}^{\infty} \frac{dw}{2\pi} \int \frac{d^3 k}{(2\pi)^3} \cdot \frac{1}{\varepsilon - w - \frac{(\vec{p} - \vec{k})^2}{2m} + i\delta}$$

poles in the upper half-plane.

$$x \frac{w_0(k)}{2} \left[\frac{1}{w - w_0(k) + i\delta} - \frac{1}{w + w_0(k) - i\delta} \right] =$$

pole in the lower h.p. upper.

$$= -2i\pi \cdot ig^2 \cdot \frac{i}{2\pi} \cdot \frac{1}{2} \frac{w_0(k)}{\varepsilon - w_0(k) - \frac{(\vec{p} - \vec{k})^2}{2m} + i\delta}$$

$$= \frac{g^2}{2} \cdot \int \frac{w_0(k)}{\varepsilon - w_0(k) - \frac{(\vec{p} - \vec{k})^2}{2m} + i\delta} \frac{d^3 k}{(2\pi)^3}$$

Analysis of the expression.

Let us analyze whether there is imaginary part and whether imaginary part is zero.

$x_{min} = (\vec{p} - \vec{k})^2$, \Rightarrow if argument of δ -function is > 0 along

then $\text{Im } \Sigma = 0$.

We have $(\vec{p} - \vec{k})^2 - 2m [\varepsilon - c\vec{k}] > 0$.

$$\vec{p}^2 + \vec{k}^2 - 2\vec{p}\cdot\vec{k} - \frac{2mc\vec{p}}{2m} + 2mc\vec{k} > 0$$

$$\vec{k} + 2(m\vec{c} - \vec{p}) > 0$$

Let \vec{p} small and > 0 if

$$(m\vec{c} - \vec{p}) > 0 \Rightarrow \text{Im } \Sigma = 0$$

which yields

$$\Sigma = \frac{g^2}{2(2\pi)^2 p} \int_0^{k_0} k dk \int_{p-k}^{p+k} \frac{ck \rho dq}{\epsilon - \frac{q^2}{2m} - ck + i0}, \quad \boxed{\frac{1}{x+i0} = \frac{1}{x} - i\pi \delta(x)}$$

Now we integrate over $\mathbb{R} = p^2$:

$$\Sigma = \frac{g^2 mc}{8\pi^2 p} \int_0^{k_0} \ln \left| \frac{\epsilon - (p-k)^2/2m - ck}{\epsilon - (p+k)^2/2m - ck} \right| \frac{k^2 dk}{k^2} - \\ - i \frac{g^2 mc}{8\pi p} \int_0^{k_0} k^2 dk \int_{(p-k)^2}^{(p+k)^2} \delta(x - 2m(\epsilon - ck)) dx$$

Analogies:

When $p < mc$, $\Rightarrow \Sigma$ is real. \Rightarrow particle is stable!

Effects coming from $p > mc$ will be discussed on Wednesday.

We are interested in Σ near the mass shell $\epsilon = \frac{p^2}{2m}$
and when p is small. $p \ll mc$, $\Delta = \left(\epsilon - \frac{p^2}{2m} \right) \ll \frac{mc^2}{2m}$, $V = \frac{p^2}{2m}$

Expanding in Δ and V we obtain

$$(p-k)^2 < 2m \left(\frac{p^2}{2m} - ck \right) < (p+k)^2 \quad k, p > 0$$

\Rightarrow equivalent to

$$0 < k < 2(p-mc)$$

~~if $p \gg mc$~~
~~then k must be small~~

In Σ is finite

$$\Sigma = \frac{g^2 m c}{8\pi^2 p} \int_0^{k_D} \ln \left| \frac{k^2/2m + (c-v)k - \Delta}{k^2/2m + (c+v)k - \Delta} \right| k^2 dk =$$

$$= \frac{g^2 m c}{8\pi^2 p} \int_0^{k_D} \left(-\frac{2vk}{ck + k^2/2m} - \frac{4\Delta v k}{2(ck + k^2/2m)^2} - \frac{2v^3 k^3}{3(ck + k^2/2m)^3} \right) \times k^2 dk$$

These three terms yield

$$\Sigma = \epsilon_0 - d_1 \Delta - d_2 \frac{p^3}{2m} \text{ form of } \Sigma. \text{ Moreover,}$$

since $C \ll \frac{k_D}{m}$, in denominators we can safely neglect k_C with respect to $k^2/2m$. This yields:

$$\epsilon_0 = -\frac{g^2 c}{4\pi^2} \int_0^{k_D} \frac{k^3 dk}{ck + k^2/2m} = -\frac{g^2 m c k_D^2}{4\pi^3}.$$

Within log accuracy we have.

$$d_1 = \frac{g^2 m^2 c}{\pi^2} \int_0^{k_D} \frac{k^3 dk}{(k^2 + 2mck)^2} = \frac{g^2 m^2 c}{\pi^2} \ln \frac{k_D}{mc} \begin{cases} G_0^{-1} - 2 = \\ = \omega - \frac{p^2}{2m} - \epsilon_0 + d_1 \frac{p^2}{2m} \\ \uparrow \omega = \frac{p^2}{2m} (1 - d_1) \end{cases}$$

$$d_2 = \frac{2g^2 m^3 c}{3\pi^2} \cdot \frac{2}{m} \int_0^{k_D} \frac{k^5 dk}{(k^2 + 2mck)^3} = \frac{4g^2 m^2 c}{3\pi^2} \ln \frac{k_D}{mc} = \frac{4}{3} d_1.$$

Dispersion relation is determined from $G_0^{-1} - 2 = 0 \Rightarrow$

$$\Rightarrow \frac{m_p}{m} = \frac{1}{1 - d_2} > 1 \rightarrow \text{electron dresser rather than unloader.}$$

4. Now let us turn to the next problem:

Cherenkov radiation of sound.

We will show that $\sum (\varepsilon, \vec{p})$ self-energy has a finite imaginary part when $v = \frac{p}{m} > c$. This is because if electron is faster than sound, it emits photons.

From (15) we see that ~~apart from integration over x~~ yields a finite value if:

$$(p-v)^2 < 2m \left(\frac{p^2}{m} - ct \right) < (p+ct)^2 \quad t, p > 0.$$

This condition is equivalent to $0 < k < 2(p-mc)$, and

Therefore $\text{Im} \Sigma$ is finite when $v = \frac{p}{m} > c$.

Probability to emit a photon in unit time is given by

$$2\delta = -2 \text{Im} \Sigma = 2 \frac{g^2 mc}{8\pi p} \int_0^{2m(v-c)} k^2 dk = \frac{2g^2 m^3}{3\pi} \frac{c}{v} \cdot (v-c)^3$$

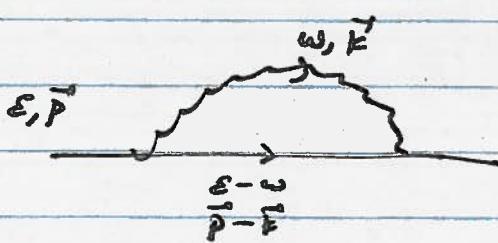
From here we see that electron emits photons when

$$v > c.$$

Probability of ~~processes~~ of emission of a photon decreases when $v \rightarrow c$ and vanishes when $v < c$.

Now let us find the angular distribution of intensity of emitted sound.

We have such a process:



Electron has momentum

$$\vec{q} = \vec{p} - \vec{k} \Rightarrow q^2 = k^2 + p^2 - 2pk \cos \theta.$$

From conservation of energy we have $\frac{p^2}{2m} = \frac{q^2}{2m} + ck$.

Therefore $\cos \theta = \frac{(k + 2mc)}{2p}$. and we can pass from $dk \rightarrow d\Omega$.

Then the imaginary part of the self-energy reads:

$$\text{Im } \Sigma = -\frac{g^2 mc}{8\pi p} \int k^2 dk = \frac{g^2 mc}{4\pi} \int_0^{\theta_{\max}} [2p \cos \theta - 2mc]^2 d\Omega =$$

$$= \frac{g^2 (mc)^3}{\pi} \int_0^{\theta_{\max}} \left(\frac{v}{c} \cos \theta - 1 \right)^2 \sin \theta d\theta$$

where $\theta_{\max} = \arccos(c/v)$ - maximal value of angle

under which a phonon ~~can be~~ emitted.

Integrand $f(\theta) = \# \cdot \left(\frac{v}{c} \cos \theta - 1 \right)^2 \cdot \sin \theta$

defines the angular distribution of intensity of emitted sound inside a cone with angle θ_{\max} .

$\delta(\omega + \omega_0(k))$ in $\text{Im} \Sigma$ can be neglected as it corresponds to absorption of a phonon. Thus we obtain:

$$\text{Im} \Sigma = -\frac{\pi^2}{2} g^2 \int \omega_0(\vec{k}) \delta\left(\varepsilon - \omega - \frac{(\vec{p} - \vec{k})^2}{2m}\right) \delta(\omega - \omega_0(k)) \frac{d^3 k}{(2\pi)^3} \frac{d\omega}{2\pi}$$

Integration over ω is straightforward and yields

$$\text{Im} \Sigma(\varepsilon, \vec{p}) = -\frac{\pi}{4} g^2 \int \omega_0(k) \delta\left(\varepsilon - \omega_0(k) - \frac{(\vec{p} - \vec{k})^2}{2m}\right) \frac{d^3 k}{(2\pi)^3}$$

Taking $\text{Im} \Sigma$ on a mass shell we obtain

$$\text{Im} \Sigma \left(\varepsilon = \frac{\vec{p}^2}{2m}, \vec{p}\right) = -\frac{\pi}{4} g^2 c \int \delta\left(\frac{k}{2m} + c - \frac{\vec{p}}{m} \cos\theta\right) \frac{k^2 dk}{(2\pi)^2} \sin\theta$$

Now we integrate over k with the help of the δ -function address which yields the same expression we had before.

This is the same Cherenkov effect but for electrons and phonons. - 16-

4. b) Second solution:

Physical meaning of $\text{Im}\Sigma$ is inverse time ~~average~~ decay into an electron and phonon.

Resultant particles, electron and phonon are real particles, which means their energies and momenta are connected through dispersion relations. Therefore, to calculate $\text{Im}\Sigma$ it is sufficient to extract the contribution coming from mass-shell of intermediate particles:

$$G_0(\varepsilon, \vec{p}) \rightarrow \text{Im } G_0(\varepsilon, \vec{p}) = -i\bar{\omega} \delta(\varepsilon - \bar{\omega}_{\text{em}}) \quad \begin{matrix} \text{emission} \\ \downarrow \text{very} \end{matrix} \quad \begin{matrix} \text{absorption} \\ \uparrow \end{matrix}$$

$$D_0(\omega, \vec{k}) \rightarrow \text{Im } D_0(\omega, \vec{k}) = -i\frac{\pi}{2} \omega_0(k) \left[\delta(\omega - \omega_0(k)) + \delta(\omega + \omega_0(k)) \right]$$

In other words the intermediate particles are not virtual. They are real and therefore

$$\text{Im}\Sigma = g^2 \int \text{Im } G_0(\varepsilon - \omega, \vec{p} - \vec{k}) \text{Im } D_0(\omega, \vec{k}) \frac{d^3 k}{(2\pi)^3} \frac{d\omega}{2\bar{\omega}}$$

But formally speaking, this is an exact mathematical identity.